## Scotogenic model for co-bimaximal mixing

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Abstract: We present a scotogenic model, i.e. a one-loop neutrino mass model with dark right-handed neutrino gauge singlets and one inert dark scalar gauge doublet $\eta$, which has symmetries that lead to co-bimaximal mixing, i.e. to an atmospheric mixing angle $\theta_{23}=45^{\circ}$ and to a $C P$-violating phase $\delta= \pm \pi / 2$, while the mixing angle $\theta_{13}$ remains arbitrary. The symmetries consist of softly broken lepton numbers $L_{\alpha}(\alpha=e, \mu, \tau)$, a non-standard $C P$ symmetry, and three $\mathbb{Z}_{2}$ symmetries. We indicate two possibilities for extending the model to the quark sector. Since the model has, besides $\eta$, three scalar gauge doublets, we perform a thorough discussion of its scalar sector. We demonstrate that it can accommodate a Standard Model-like scalar with mass 125 GeV , with all the other charged and neutral scalars having much higher masses.

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## 1 Introduction

With the experimental finding that the lepton mixing angle $\theta_{13}$ is nonzero, many theoretical neutrino mass models fell into disfavour. An exception is the model in ref. [1], in which there is a relation

$$
S \mathcal{M}_{\nu} S=\mathcal{M}_{\nu}^{*}, \quad \text { where } S=\left(\begin{array}{lll}
1 & 0 & 0  \tag{1.1}\\
0 & 0 & 1 \\
0 & 1 & 0
\end{array}\right)
$$

and $\mathcal{M}_{\nu}$ is the (symmetric) light-neutrino Majorana mass matrix in the basis where the charged-lepton mass matrix is diag $\left(m_{e}, m_{\mu}, m_{\tau}\right)$. The condition (1.1) leads to $\sin ^{2} \theta_{23}=1 / 2$ and (provided $\left.\sin \theta_{13} \neq 0\right) \cos \delta=0$, which is in agreement with the phenomenology [2-5]; this situation has recently been dubbed 'co-bimaximal mixing' [6]. ${ }^{1}$ A relevant point is that the condition (1.1) does not restrict the neutrino masses; it only restricts lepton mixing. Actually, as a consequence of the condition (1.1), the lepton mixing matrix has the form [1]

$$
U=\left(\begin{array}{lll}
u_{1} \eta_{1} & u_{2} \eta_{2} & u_{3} \eta_{3}  \tag{1.2}\\
w_{1} \eta_{1} & w_{2} \eta_{2} & w_{3} \eta_{3} \\
w_{1}^{*} \eta_{1} & w_{2}^{*} \eta_{2} & w_{3}^{*} \eta_{3}
\end{array}\right)
$$

with $u_{j} \geq 0,\left|w_{j}\right|^{2}=\left(1-u_{j}^{2}\right) / 2$, and $\eta_{j}^{2}= \pm 1$ for $j=1,2,3$. It is clear from equation (1.2) that $\left|U_{\mu j}\right|=\left|U_{\tau j}\right| \forall j=1,2,3[9-11]$. Note that the condition (1.1) entails three restrictions on lepton mixing:

[^0]1. The atmospheric mixing angle $\theta_{23}$ is maximal, i.e. $\sin \left(2 \theta_{23}\right)=1$.
2. The $C P$-violating phase $\delta$ is $\pm \pi / 2$.
3. The Majorana phase factors in effective neutrinoless $\beta \beta$ decay are $\pm 1$.

Because the predictions of condition (1.1) do not depend on the neutrino masses, it is possible that, in some multi-Higgs-doublet models, co-bimaximal mixing is not disturbed by the one-loop corrections to the neutrino mass matrix [12]. This may, in particular, be the case in a 'scotogenic' model [13]. In such a model, the masses of the light neutrinos have radiative origin and the particles in the loop that generates them belong to the dark sector of astrophysics, i.e. they are prevented from mixing with the ordinary particles by an unbroken (usually $\mathbb{Z}_{2}$ ) symmetry.

The purpose of this paper is to propose a scotogenic model for the neutrino masses which predicts co-bimaximal mixing. ${ }^{2}$ This is done in section 2 . In section 3 we expose two possible extensions of that model to the quark sector. An analysis of the scalar potential of the model and of its compatibility with the recently discovered scalar of mass 125 GeV is performed in section 4. We summarize our findings in section 5. Appendix A collects some formulae from ref. [12] which are used in section 2 .

## 2 The model for the lepton sector

Our model is an extension of the Standard Model with gauge symmetry $\mathrm{SU}(2) \times \mathrm{U}(1)$. The usual fermion multiplets are three $D_{\alpha L}$ and three $\alpha_{R}(\alpha=e, \mu, \tau)$. Besides, we introduce three right-handed neutrinos $\nu_{\alpha R}$; they belong to the dark sector of the model. Our model has four scalar doublets:

$$
\begin{equation*}
\phi_{j}=\binom{\phi_{j}^{+}}{\phi_{j}^{0}} \quad(j=1,2,3) \quad \text { and } \quad \phi_{4} \equiv \eta=\binom{\eta^{+}}{\eta^{0}} \tag{2.1}
\end{equation*}
$$

The doublet $\phi_{1}$ gives mass to the electron, $\phi_{2}$ gives mass to the muon, and $\phi_{3}$ gives mass to the $\tau$ lepton; the doublet $\phi_{4} \equiv \eta$ belongs to the dark sector. We shall also use the conjugate doublets $\tilde{\phi}_{j}=\left(\phi_{j}^{0^{*}},-\phi_{j}^{-}\right)^{T}$ and $\tilde{\eta}=\left(\eta^{0^{*}},-\eta^{-}\right)^{T}$.

The symmetries of our model are the following:

- $\mathbb{Z}_{2}^{\text {(dark })}: \eta \rightarrow-\eta, \nu_{e R} \rightarrow-\nu_{e R}, \nu_{\mu R} \rightarrow-\nu_{\mu R}$, and $\nu_{\tau R} \rightarrow-\nu_{\tau R}$. This is an exact symmetry that prevents dark matter from mixing with ordinary matter. It is broken neither softly nor spontaneously, because the vacuum expectation value (VEV) of $\eta$ is zero. ${ }^{3}$

[^1]- The flavour lepton numbers $L_{\alpha}$. They are broken only softly by the Majorana mass terms of the $\nu_{\alpha R}$ :

$$
\mathcal{L}_{\text {Majorana }}=-\frac{1}{2}\left(\overline{\nu_{e R}}, \overline{\nu_{\mu R}}, \overline{\nu_{\tau R}}\right) M_{R} C\left(\begin{array}{l}
{\overline{\nu_{e R}}}^{T}  \tag{2.2}\\
{\overline{\nu_{\mu R}}}^{T} \\
{\overline{\nu_{\tau R}}}^{T}
\end{array}\right)+\text { H.c., }
$$

where $C$ is the charge-conjugation matrix in Dirac space and $M_{R}$ is a symmetric matrix in flavour space.

- $\mathbb{Z}_{2}^{(1)}: \phi_{1} \rightarrow-\phi_{1}, e_{R} \rightarrow-e_{R}, \mathbb{Z}_{2}^{(2)}: \phi_{2} \rightarrow-\phi_{2}, \mu_{R} \rightarrow-\mu_{R}$, and $\mathbb{Z}_{2}^{(3)}: \phi_{3} \rightarrow-\phi_{3}$, $\tau_{R} \rightarrow-\tau_{R}$. Because of these symmetries and of the $L_{\alpha}$, the lepton Yukawa Lagrangian is

$$
\begin{align*}
\mathcal{L}_{\ell \text { Yukawa }}= & -y_{1} \overline{\nu_{e R}} \tilde{\eta}^{\dagger} D_{e L}-y_{2} \overline{\nu_{\mu R}} \tilde{\eta}^{\dagger} D_{\mu L}-y_{3} \overline{\nu_{\tau R}} \tilde{\eta}^{\dagger} D_{\tau L}  \tag{2.3a}\\
& -y_{4} \overline{e_{R}} \phi_{1}^{\dagger} D_{e L}-y_{5} \overline{\mu_{R}} \phi_{2}^{\dagger} D_{\mu L}-y_{6} \overline{\tau_{R}} \phi_{3}^{\dagger} D_{\tau L}+\text { H.c. } \tag{2.3b}
\end{align*}
$$

The $\mathbb{Z}_{2}^{(j)}(j=1,2,3)$ are broken spontaneously, through the VEVs $\langle 0| \phi_{j}^{0}|0\rangle=v_{j} / \sqrt{2}$, to give mass to the charged leptons:

$$
\begin{equation*}
m_{e}=\left|\frac{y_{4} v_{1}}{\sqrt{2}}\right|, \quad m_{\mu}=\left|\frac{y_{5} v_{2}}{\sqrt{2}}\right|, \quad m_{\tau}=\left|\frac{y_{6} v_{3}}{\sqrt{2}}\right| \tag{2.4}
\end{equation*}
$$

Besides, the $\mathbb{Z}_{2}^{(j)}$ are also broken softly ${ }^{4}$ through quadratic terms in the scalar potential.

- The $C P$ symmetry

$$
C P:\left\{\begin{array}{rl}
D_{L} & \rightarrow i \gamma_{0} C S{\overline{D_{L}}}^{T},  \tag{2.5}\\
\ell_{R} & \rightarrow i \gamma_{0} C S{\overline{\ell_{R}}}^{T}, \\
\nu_{R} & \rightarrow i \gamma_{0} C S{\overline{\nu_{R}}}^{T}, \\
\phi & \rightarrow S \phi^{*}, \\
\eta & \rightarrow \eta^{*},
\end{array} \quad D_{L}=\left(\begin{array}{c}
D_{e L} \\
D_{\mu L} \\
D_{\tau L}
\end{array}\right), \quad \ell_{R}=\left(\begin{array}{c}
e_{R} \\
\mu_{R} \\
\tau_{R}
\end{array}\right),\right.
$$

Because of this symmetry, in equation (2.2)

$$
M_{R}=\left(\begin{array}{ccc}
x & y & y^{*}  \tag{2.6}\\
y & z & w \\
y^{*} & w & z^{*}
\end{array}\right)
$$

with real $x$ and $w$, i.e. $S M_{R} S=M_{R}^{*}$; moreover, in equation (2.3) $y_{1}$ and $y_{4}$ are real and $y_{3}=y_{2}^{*}, y_{6}=y_{5}^{*}$. Therefore,

$$
\begin{equation*}
\frac{m_{\mu}}{m_{\tau}}=\left|\frac{v_{2}}{v_{3}}\right| \tag{2.7}
\end{equation*}
$$

[^2]i.e. the small ratio of muon to $\tau$-lepton mass is explained through a small ratio of VEVs [16]. The symmetry $C P$ is not broken softly ${ }^{5}$ but it is broken spontaneously through the VEVs $v_{j}$, especially through $\left|v_{2}\right| \neq\left|v_{3}\right| .{ }^{6}$

As compared to the model in ref. [1], the present model has an extra doublet $\eta$, whose vanishing VEV causes neutrino mass generation to occur only at the one-loop level. However, as we will show below, the very same mechanism that produces co-bimaximal mixing at the tree level in the model of ref. [1] is effective at the one-loop level in the model of this paper.

In our model, just as in the original model of Ma [13], dark matter may be either spin-one half - the lightest particle arising from the mixture of $\nu_{e R}, \nu_{\mu R}$, and $\nu_{\tau R}$ - or spin-zero - the lightest of the two components $\varphi_{1,2}$ of $\eta^{0}$ - depending on which of them is lighter. No other fields are needed in principle to account for the dark matter.

In the scalar potential, a crucial role is played by the $C P$-invariant terms

$$
\begin{equation*}
\xi_{1}\left[\left(\phi_{1}^{\dagger} \eta\right)^{2}+\left(\eta^{\dagger} \phi_{1}\right)^{2}\right]+\xi_{2}\left[\left(\phi_{2}^{\dagger} \eta\right)^{2}+\left(\eta^{\dagger} \phi_{3}\right)^{2}\right]+\xi_{3}\left[\left(\phi_{3}^{\dagger} \eta\right)^{2}+\left(\eta^{\dagger} \phi_{2}\right)^{2}\right] \tag{2.8}
\end{equation*}
$$

where $\xi_{1}=\xi_{1}^{*}$ and $\xi_{3}=\xi_{2}^{*}$ because of Hermiticity. Let us write

$$
\begin{equation*}
\phi_{4}^{0} \equiv \eta^{0}=e^{i \gamma} \frac{\varphi_{1}+i \varphi_{2}}{\sqrt{2}} \tag{2.9}
\end{equation*}
$$

where the fields $\varphi_{1}$ and $\varphi_{2}$ are real and the phase $\gamma$ is defined such that

$$
\begin{equation*}
\mu^{2} \equiv e^{2 i \gamma} \sum_{j=1}^{3} \xi_{j} \frac{v_{j}^{* 2}}{2} \tag{2.10}
\end{equation*}
$$

is real and positive. Then, the terms (2.8) generate a mass term

$$
\begin{equation*}
\mu^{2}\left(\varphi_{1}^{2}-\varphi_{2}^{2}\right) \tag{2.11}
\end{equation*}
$$

which means that $\varphi_{1}$ and $\varphi_{2}$ are mass eigenfields with distinct masses. The term (2.11) is the only one that makes the masses of $\varphi_{1}$ and $\varphi_{2}$ different; all other terms in the scalar potential contain $\left|\eta^{0}\right|^{2}=\left(\varphi_{1}^{2}+\varphi_{2}^{2}\right) / 2$.

Now we make use of the results in appendix A. In the notation of equation (A.1), equation (2.3a) means that $\Delta_{1}=\Delta_{2}=\Delta_{3}=0$ and $\Delta_{4}=\operatorname{diag}\left(y_{1}, y_{2}, y_{2}^{*}\right)$; notice that $S \Delta_{4} S=\Delta_{4}^{*}$. In the notation of equation (A.2), equation (2.9) reads $\mathcal{V}_{4 \varphi_{1}}=e^{i \gamma}$ and $\mathcal{V}_{4 \varphi_{2}}=i e^{i \gamma}$. Then, according to equation (A.3), $\Delta_{\varphi_{1}}=e^{i \gamma} \Delta_{4}$ and $\Delta_{\varphi_{2}}=i e^{i \gamma} \Delta_{4}$. Applying equation (A.4) we find the one-loop contribution to $\mathcal{M}_{\nu}$ :

$$
\begin{equation*}
\delta \mathcal{M}_{\nu}=\frac{e^{2 i \gamma}}{32 \pi^{2}}\left[\Delta_{4} W^{*}\left(\frac{m_{\varphi_{1}}^{2}}{\widetilde{m}} \ln \frac{\widetilde{m}^{2}}{m_{\varphi_{1}}^{2}}\right) W^{\dagger} \Delta_{4}-\Delta_{4} W^{*}\left(\frac{m_{\varphi_{2}}^{2}}{\widetilde{m}} \ln \frac{\widetilde{m}^{2}}{m_{\varphi_{2}}^{2}}\right) W^{\dagger} \Delta_{4}\right] \tag{2.12}
\end{equation*}
$$

[^3]where the matrices $W$ and $\widetilde{m}$ are defined through equation (A.5). Note that there is no contribution to $\delta \mathcal{M}_{\nu}$ from a loop with $Z^{0}$ because the VEV of $\eta$ is assumed to vanish; therefore, the Dirac neutrino mass matrix $M_{D}$ in line (A.4b) also vanishes.

In the limit $\mu^{2} \rightarrow 0$, the masses of $\varphi_{1}$ and $\varphi_{2}$ become equal and the contributions of $\varphi_{1}$ and $\varphi_{2}$ to $\delta \mathcal{M}_{\nu}$ exactly cancel each other; the light neutrinos then remain massless at the one-loop level [13]. This happens in the limit where all the terms in equation (2.8) vanish. Indeed, in that limit the full Lagrangian is invariant under the $\mathrm{U}(1)$ symmetry

$$
\begin{equation*}
D_{L} \rightarrow e^{i \psi} D_{L}, \quad \ell_{R} \rightarrow e^{i \psi} \ell_{R}, \quad \eta \rightarrow e^{-i \psi} \eta, \tag{2.13}
\end{equation*}
$$

which forbids light-neutrino masses [13]. We remark that there are, in the scotogenic model of this paper, several mechanisms for potentially suppressing the light-neutrino masses, viz.

- a large seesaw scale, i.e. large heavy-neutrino masses in $\widetilde{m}$;
- small Yukawa couplings of $\nu_{R}$, i.e. small $\Delta_{4}$;
- small couplings $\xi_{1,2,3}$ in equation (2.8), hence $m_{\varphi_{1}}$ and $m_{\varphi_{2}}$ very close to each other, because of an approximate symmetry (2.13);
- the $\left(32 \pi^{2}\right)^{-1}$ factor in equation (2.12) from the loop integral.

Let us present a benchmark for all these suppressing factors. Let both $\xi_{1,2,3}$ and $y_{1,2}$ be of order $10^{-2}$. With $\left|v_{1,2,3}\right| \sim 100 \mathrm{GeV}$ one then obtains $\left|m_{\varphi_{1}}-m_{\varphi_{2}}\right| \sim 10 \mathrm{GeV}$. Assuming $m_{\varphi_{1,2}} \sim 100 \mathrm{GeV}$, one requires $\tilde{m} \sim 10^{7-8} \mathrm{GeV}$ in order to obtain $\delta \mathcal{M}_{\nu} \sim 0.1 \mathrm{eV}$. One concludes that the main suppression still originates in the high seesaw scale. However, with small $\xi_{1,2,3}$ and $y_{1,2}$, of order $10^{-3}$ or $10^{-4}$, the seesaw scale could easily be in the TeV range and thus accessible to the LHC.

Next we exploit the $C P$-invariance properties, viz. $S \Delta_{4} S=\Delta_{4}^{*}$ and $S M_{R} S=M_{R}^{*}$. Equation (2.12) may be rewritten

$$
\begin{equation*}
e^{-2 i \gamma} \delta \mathcal{M}_{\nu}=\Delta_{4} W^{*} \hat{A} W^{\dagger} \Delta_{4}, \quad \text { with } \hat{A}=\frac{1}{32 \pi^{2}}\left(\frac{m_{\varphi_{1}}^{2}}{\tilde{m}} \ln \frac{\widetilde{m}^{2}}{m_{\varphi_{1}}^{2}}-\frac{m_{\varphi_{2}}^{2}}{\widetilde{m}} \ln \frac{\widetilde{m}^{2}}{m_{\varphi_{2}}^{2}}\right) . \tag{2.14}
\end{equation*}
$$

Now,

$$
\begin{equation*}
S M_{R} S=M_{R}^{*} \Rightarrow\left(W^{\dagger} S W^{*}\right)^{*} \widetilde{m}=\widetilde{m}\left(W^{\dagger} S W^{*}\right) \Rightarrow W^{\dagger} S W^{*}=X, \tag{2.15}
\end{equation*}
$$

where $X$ is a diagonal sign matrix [1]. This is because, according to the assumptions of the seesaw mechanism, all the diagonal matrix elements of $\widetilde{m}$, i.e. all the heavy-neutrino masses, are nonzero. Using equation (2.15) we derive

$$
\begin{align*}
S\left(\Delta_{4} W^{*} \hat{A} W^{\dagger} \Delta_{4}\right) S & =\left(S \Delta_{4} S\right)\left(S W^{*}\right) \hat{A}\left(W^{\dagger} S\right)\left(S \Delta_{4} S\right) \\
& =\Delta_{4}^{*} W X \hat{A} X W^{T} \Delta_{4}^{*} \\
& =\Delta_{4}^{*} W \hat{A} W^{T} \Delta_{4}^{*} \\
& =\left(\Delta_{4} W^{*} \hat{A} W^{\dagger} \Delta_{4}\right)^{*}, \tag{2.16}
\end{align*}
$$

i.e.

$$
\begin{equation*}
S\left(e^{-2 i \gamma} \delta \mathcal{M}_{\nu}\right) S=\left(e^{-2 i \gamma} \delta \mathcal{M}_{\nu}\right)^{*} \tag{2.17}
\end{equation*}
$$

Thus, after a physically meaningless rephasing, $\delta \mathcal{M}_{\nu}$ displays the defining feature (1.1) of co-bimaximal mixing.

### 2.1 Approximation to the Higgs boson

We use the notation of equation (A.2). The matrix $\mathcal{V}$ is complex $4 \times 8$ and, according to ref. [19],

$$
\begin{equation*}
\tilde{\mathcal{V}}=\binom{\Re \mathcal{V}}{\Im \mathcal{V}} \tag{2.18}
\end{equation*}
$$

is $8 \times 8$ orthogonal. The last row of $\mathcal{V}$ corresponds to $\phi_{4}^{0} \equiv \eta^{0}$. For definiteness, we let the last two columns of $\mathcal{V}$ correspond to $\varphi_{1}$ and $\varphi_{2}$, which belong to the dark sector and do not mix with all the other scalars. Therefore, for practical purposes $\mathcal{V}$ is just a $3 \times 6$ matrix. By definition, $S_{1}^{0}=G^{0}$ is the Goldstone boson and [19]

$$
\begin{equation*}
\mathcal{V}_{j 1}=i \frac{v_{j}}{v} \quad \text { for } j=1,2,3, \quad \text { where } v \equiv \sqrt{\sum_{j=1}^{3}\left|v_{j}\right|^{2}} \tag{2.19}
\end{equation*}
$$

The couplings of $S_{b}^{0}(b=2, \ldots, 6)$ to the gauge bosons are given by [19]

$$
\begin{equation*}
\frac{g}{v}\left(m_{W} W_{\mu}^{+} W^{\mu-}+\frac{m_{Z} Z_{\mu} Z^{\mu}}{2 c_{w}}\right) \sum_{b=2}^{8} S_{b}^{0} \Re\left(\sum_{j=1}^{3} v_{j}^{*} \mathcal{V}_{j b}\right) \tag{2.20}
\end{equation*}
$$

Therefore, a given $S_{b}^{0}$ couples to the gauge bosons with exactly the same strength as the Higgs boson of the Standard Model if

$$
\begin{equation*}
g_{S V V} \equiv \Re\left(\sum_{j=1}^{3} \frac{v_{j}^{*}}{v} \mathcal{V}_{j b}\right)=1 \tag{2.21}
\end{equation*}
$$

Notice that, because both three-vectors $\left(v_{1}^{*} / v, v_{2}^{*} / v, v_{3}^{*} / v\right)$ and $\left(\mathcal{V}_{1 b}, \mathcal{V}_{2 b}, \mathcal{V}_{3 b}\right)$ have unit modulus, $\left|g_{S V V}\right| \leq 1$. Therefore, equation (2.21) holds in a limit situation.

According to equation (2.3b), the scalars $S_{b}^{0}$ couple to the $\tau$ lepton through

$$
\begin{equation*}
-\frac{1}{\sqrt{2}} \sum_{b=1}^{6} S_{b}^{0} \bar{\tau}\left(\mathcal{V}_{3 b} y_{6}^{*} \gamma_{R}+\mathcal{V}_{3 b}^{*} y_{6} \gamma_{L}\right) \tau=-m_{\tau} \sum_{b=1}^{6} S_{b}^{0} \bar{\tau}\left(\frac{\mathcal{V}_{3 b}}{v_{3}} \gamma_{R}+\frac{\mathcal{V}_{3 b}^{*}}{v_{3}^{*}} \gamma_{L}\right) \tau \tag{2.22}
\end{equation*}
$$

where $\gamma_{R, L}$ are the projectors of chirality in Dirac space. In equation (2.22) we have assumed, without loss of generality, $y_{6} v_{3}^{*}$ to be real and positive. Therefore, a given $S_{b}^{0}$ couples to the $\tau$ lepton in the same way as the Higgs boson if

$$
\begin{equation*}
\frac{v \mathcal{V}_{3 b}}{v_{3}}=1 \tag{2.23}
\end{equation*}
$$

## 3 Extension of the model to the quark sector

It is non-trivial to extend our model to the quark sector because the $C P$ symmetry relates the Yukawa couplings of $\phi_{2}$ to those of $\phi_{3}$; moreover, some quarks must couple to $\phi_{2}$ and, correspondingly, other quarks must couple to $\phi_{3}$ - in order that $C P$ violation, which is generated through $v_{1}^{*} v_{2} \neq v_{1} v_{3}^{*}$, manifests itself in the CKM matrix $V$.

We firstly expound some notation. The quark Yukawa Lagrangian is

$$
\mathcal{L}_{\text {quark Yukawa }}=-\left(\overline{Q_{L 1}}, \overline{Q_{L 2}}, \overline{Q_{L 3}}\right) \sum_{j=1}^{3}\left[\phi_{j} \Gamma_{j}\left(\begin{array}{c}
n_{R 1}  \tag{3.1}\\
n_{R 2} \\
n_{R 3}
\end{array}\right)+\tilde{\phi}_{j} \Delta_{j}\left(\begin{array}{c}
p_{R 1} \\
p_{R 2} \\
p_{R 3}
\end{array}\right)\right]+\text { H.c., }
$$

where $\overline{Q_{L j}}=\left(\overline{p_{L j}}, \overline{n_{L j}}\right)$ for $j=1,2,3$. The mass matrices are $M_{n}=\sum_{j=1}^{3}\left(v_{j} / \sqrt{2}\right) \Gamma_{j}$ and $M_{p}=\sum_{j=1}^{3}\left(v_{j}^{*} / \sqrt{2}\right) \Delta_{j}$. They are diagonalized as

$$
\begin{align*}
U_{L}^{n \dagger} M_{n} U_{R}^{n} & =\operatorname{diag}\left(m_{d}, m_{s}, m_{b}\right) \equiv M_{d}  \tag{3.2a}\\
U_{L}^{p \dagger} M_{p} U_{R}^{p} & =\operatorname{diag}\left(m_{u}, m_{c}, m_{t}\right) \equiv M_{u} \tag{3.2b}
\end{align*}
$$

where the matrices $U_{L, R}^{n, p}$ are unitary. The physical quarks are given by

$$
\left(\begin{array}{c}
n_{R 1}  \tag{3.3}\\
n_{R 2} \\
n_{R 3}
\end{array}\right)=U_{R}^{n}\left(\begin{array}{c}
c_{R} \\
s_{R} \\
b_{R}
\end{array}\right), \quad\left(\begin{array}{c}
p_{R 1} \\
p_{R 2} \\
p_{R 3}
\end{array}\right)=U_{R}^{p}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right),
$$

and analogously for the left-handed fields. The quark mixing matrix is $V=U_{L}^{p \dagger} U_{L}^{n}$.

### 3.1 Extension 1

One may include the quarks in the symmetries $\mathbb{Z}_{2}^{(j)}$ as follows:

$$
\begin{array}{ll}
\mathbb{Z}_{2}^{(1)}: & \phi_{1}, e_{R}, \text { and } Q_{L 1} \text { change sign; } \\
\mathbb{Z}_{2}^{(2)}: & \phi_{2}, \mu_{R}, \text { and } Q_{L 2} \text { change sign; } \\
\mathbb{Z}_{2}^{(3)}: & \phi_{3}, \tau_{R}, \text { and } Q_{L 3} \text { change sign. } \tag{3.4c}
\end{array}
$$

Then,

$$
\begin{align*}
\Gamma_{1} & =\left(\begin{array}{c}
R_{1} \\
0_{1 \times 3} \\
0_{1 \times 3}
\end{array}\right),  \tag{3.5a}\\
\Gamma_{2}=\left(\begin{array}{c}
0_{1 \times 3} \\
R_{2} \\
0_{1 \times 3}
\end{array}\right), & \Gamma_{3}=\left(\begin{array}{c}
0_{1 \times 3} \\
0_{1 \times 3} \\
R_{3}
\end{array}\right),  \tag{3.5b}\\
\Delta_{1}=\left(\begin{array}{c}
R_{1}^{\prime} \\
0_{1 \times 3} \\
0_{1 \times 3}
\end{array}\right), & \Delta_{2}=\left(\begin{array}{c}
0_{1 \times 3} \\
R_{2}^{\prime} \\
0_{1 \times 3}
\end{array}\right),
\end{align*} \quad \Delta_{3}=\left(\begin{array}{c}
0_{1 \times 3} \\
0_{1 \times 3} \\
R_{3}^{\prime}
\end{array}\right), ~ \$, ~
$$

where $R_{1,2,3}$ and $R_{1,2,3}^{\prime}$ are $1 \times 3$ row matrices. Notice that both $\Gamma_{1,2,3}$ and $\Delta_{1,2,3}$ are in this extension rank 1 matrices. The quark mass matrices are

$$
M_{n}=\frac{1}{\sqrt{2}}\left(\begin{array}{l}
R_{1} v_{1}  \tag{3.6}\\
R_{2} v_{2} \\
R_{3} v_{3}
\end{array}\right), \quad M_{p}=\frac{1}{\sqrt{2}}\left(\begin{array}{l}
R_{1}^{\prime} v_{1}^{*} \\
R_{2}^{\prime} v_{2}^{*} \\
R_{3}^{\prime} v_{3}^{*}
\end{array}\right) .
$$

We define the $1 \times 3$ matrices

$$
\begin{equation*}
\bar{R}_{j} \equiv R_{j} U_{R}^{n}, \quad \bar{R}_{j}^{\prime} \equiv R_{j}^{\prime} U_{R}^{p} \tag{3.7}
\end{equation*}
$$

Writing

$$
U_{L}^{n}=\left(\begin{array}{l}
\hat{R}_{1}  \tag{3.8}\\
\hat{R}_{2} \\
\hat{R}_{3}
\end{array}\right), \quad U_{L}^{p}=\left(\begin{array}{c}
\hat{R}_{1}^{\prime} \\
\hat{R}_{2}^{\prime} \\
\hat{R}_{3}^{\prime}
\end{array}\right),
$$

where the $\hat{R}_{j}$ and $\hat{R}_{j}^{\prime}$ are $1 \times 3$ row matrices, ${ }^{7}$ one has, from equations (3.2), (3.6), and (3.7),

$$
\begin{equation*}
\bar{R}_{j} \frac{v_{j}}{\sqrt{2}}=\hat{R}_{j} M_{d}, \quad \bar{R}_{j}^{\prime} \frac{v_{j}^{*}}{\sqrt{2}}=\hat{R}_{j}^{\prime} M_{u} . \tag{3.9}
\end{equation*}
$$

(We do not use the summation convention.) The Yukawa couplings of the neutral scalars - see equation (A.2) - are

$$
\begin{align*}
&- \frac{1}{\sqrt{2}} \sum_{b=1}^{8} S_{b}^{0} \sum_{j=1}^{3}\left[\overline{n_{L j}} \mathcal{V}_{j b} R_{j}\left(\begin{array}{c}
n_{R 1} \\
n_{R 2} \\
n_{R 3}
\end{array}\right)+\overline{p_{L j}} \mathcal{V}_{j b}^{*} R_{j}^{\prime}\left(\begin{array}{c}
p_{R 1} \\
p_{R 2} \\
p_{R 3}
\end{array}\right)\right] \\
&=-\frac{1}{\sqrt{2}} \sum_{b=1}^{8} S_{b}^{0} \sum_{j=1}^{3}\left[\overline{n_{L j}} \mathcal{V}_{j b} \bar{R}_{j}\left(\begin{array}{c}
d_{R} \\
s_{R} \\
b_{R}
\end{array}\right)+\overline{p_{L j}} \mathcal{V}_{j b}^{*} \bar{R}_{j}^{\prime}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right)\right] \\
&=-\sum_{b=1}^{8} S_{b}^{0} \sum_{j=1}^{3}\left[\overline{n_{L j}} \frac{\mathcal{V}_{j b}}{v_{j}} \hat{R}_{j} M_{d}\left(\begin{array}{c}
d_{R} \\
s_{R} \\
b_{R}
\end{array}\right)+\overline{p_{L j}} \frac{\mathcal{V}_{j b}^{*}}{v_{j}^{*}} \hat{R}_{j}^{\prime} M_{u}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right)\right] \\
&=- \sum_{b=1}^{8} S_{b}^{0}\left[\left(\overline{d_{L}}, \overline{s_{L}}, \overline{b_{L}}\right) \sum_{j=1}^{3} \frac{\mathcal{V}_{j b}}{v_{j}} \hat{R}_{j}^{\dagger} \hat{R}_{j} M_{d}\left(\begin{array}{c}
d_{R} \\
s_{R} \\
b_{R}
\end{array}\right)\right. \\
&\left.\quad+\left(\overline{u_{L}}, \overline{c_{L}}, \overline{t_{L}}\right) \sum_{j=1}^{3} \frac{\mathcal{V}_{j b}^{*}}{v_{j}^{*}} \hat{R}_{j}^{\prime \dagger} \hat{R}_{j}^{\prime} M_{u}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right)\right] . \tag{3.10}
\end{align*}
$$

Defining the Hermitian matrices

$$
\begin{equation*}
H_{j} \equiv \hat{R}_{j}^{\dagger} \hat{R}_{j}=\frac{\left|v_{j}\right|^{2}}{2} M_{d}^{-1} \bar{R}_{j}^{\dagger} \bar{R}_{j} M_{d}^{-1}, \quad H_{j}^{\prime} \equiv \hat{R}_{j}^{\prime \dagger} \hat{R}_{j}^{\prime}=\frac{\left|v_{j}\right|^{2}}{2} M_{u}^{-1} \bar{R}_{j}^{\prime \dagger} \bar{R}_{j}^{\prime} M_{u}^{-1}, \tag{3.11}
\end{equation*}
$$

[^4]the Yukawa couplings of a given $S_{b}^{0}$ to the third-generation quarks are given by
\[

$$
\begin{equation*}
-S_{b}^{0} \sum_{j=1}^{3}\left[m_{b}\left(H_{j}\right)_{33} \bar{b}\left(\frac{\mathcal{V}_{j b}}{v_{j}} \gamma_{R}+\frac{\mathcal{V}_{j b}^{*}}{v_{j}^{*}} \gamma_{L}\right) b+m_{t}\left(H_{j}^{\prime}\right)_{33} \bar{t}\left(\frac{\mathcal{V}_{j b}}{v_{j}} \gamma_{L}+\frac{\mathcal{V}_{j b}^{*}}{v_{j}^{*}} \gamma_{R}\right) t\right] \tag{3.12}
\end{equation*}
$$

\]

Thus, $S_{b}^{0}$ couples to the third-generation quarks in the same way as the Higgs boson if

$$
\begin{equation*}
\sum_{j=1}^{3}\left(H_{j}\right)_{33} \frac{\mathcal{V}_{j b}}{v_{j}}=\sum_{j=1}^{3}\left(H_{j}^{\prime}\right)_{33} \frac{\mathcal{V}_{j b}}{v_{j}}=\frac{1}{v} \tag{3.13}
\end{equation*}
$$

We have not yet specified the way in which the $C P$ symmetry is to be extended to the quark sector. This may be chosen to be

$$
C P: \begin{cases}Q_{L} \rightarrow i \gamma_{0} C S{\overline{Q_{L}}}^{T}, & Q_{L}=\left(Q_{L 1}, Q_{L 2}, Q_{L 3}\right)^{T}  \tag{3.14}\\ n_{R} \rightarrow i \gamma_{0} C{\overline{n_{R}}}^{T}, \\ p_{R} \rightarrow i \gamma_{0} C{\overline{p_{R}}}^{T}, & \text { where } \\ p_{R}=\left(p_{R 1}, p_{R 2}, p_{R 3}\right)^{T} \\ & n_{R}=\left(n_{R 1}, n_{R 2}, n_{R 3}\right)^{T}\end{cases}
$$

The $C P$ symmetry (3.14) enforces real $R_{1}$ and $R_{1}^{\prime}$ and

$$
\begin{equation*}
R_{3}=R_{2}^{*}, \quad R_{3}^{\prime}=R_{2}^{\prime *} \tag{3.15}
\end{equation*}
$$

### 3.2 Extension 2

The extension of our model to the quark sector expounded in the previous subsection treats the down-type and up-type quarks in similar fashion. It possesses flavour-changing neutral Yukawa interactions (FCNYI) in both quark sectors. In this subsection we suggest a different extension, in which FCNYI are restricted to the up-type-quark sector.

Let the quarks be included in the symmetries $\mathbb{Z}_{2}^{(j)}$ as

$$
\begin{array}{ll}
\mathbb{Z}_{2}^{(1)}: & \phi_{1}, e_{R}, p_{R 1}, n_{R 1}, n_{R 2}, \text { and } n_{R 3} \text { change sign; } \\
\mathbb{Z}_{2}^{(2)}: & \phi_{2}, \mu_{R}, \text { and } p_{R 2} \text { change sign; } \\
\mathbb{Z}_{2}^{(3)}: & \phi_{3}, \tau_{R}, \text { and } p_{R 3} \text { change sign. } \tag{3.16c}
\end{array}
$$

With this extension, the Yukawa-coupling matrices $\Gamma_{2}$ and $\Gamma_{3}$ vanish outright. In extension 2, as distinct from extension 1 , the matrices $\Gamma_{2,3}$ are rank 0 while $\Gamma_{1}$ is rank 3 . Without loss of generality, one may rotate the $Q_{L}$ and the $n_{R}$ so that $v_{1} \Gamma_{1} / \sqrt{2}=M_{n}$ is equal to $M_{d}$ from the outset. Then, $U_{L}^{n}=U_{R}^{n}=\mathbb{1}_{3 \times 3}$ and the CKM matrix $V=U_{L}^{p \dagger}$.

Analogously to equation (2.22), the couplings of the neutral scalars to the down-type quarks are given by

$$
\begin{equation*}
-\sum_{b=1}^{6} S_{b}^{0} \sum_{q=d, s, b} m_{q} \bar{q}\left(\frac{\mathcal{V}_{1 b}}{v_{1}} \gamma_{R}+\frac{\mathcal{V}_{1 b}^{*}}{v_{1}^{*}} \gamma_{L}\right) q \tag{3.17}
\end{equation*}
$$

A given $S_{b}^{0}$ couples to the bottom quark in the same way as the Higgs boson if

$$
\begin{equation*}
\frac{v \mathcal{V}_{1 b}}{v_{1}}=1 \tag{3.18}
\end{equation*}
$$

Now,

$$
\begin{align*}
& \Delta_{1} \sim\left(C_{1}, 0_{3 \times 1}, 0_{3 \times 1}\right),  \tag{3.19a}\\
& \Delta_{2} \sim\left(0_{3 \times 1}, C_{2}, 0_{3 \times 1}\right),  \tag{3.19b}\\
& \Delta_{3} \sim\left(0_{3 \times 1}, 0_{3 \times 1}, C_{3}\right), \tag{3.19c}
\end{align*}
$$

where $C_{1,2,3}$ are $3 \times 1$ column vectors. The up-type-quark mass matrix is

$$
\begin{equation*}
M_{p}=\frac{1}{\sqrt{2}}\left(C_{1} v_{1}^{*}, C_{2} v_{2}^{*}, C_{3} v_{3}^{*}\right) \tag{3.20}
\end{equation*}
$$

We define the $3 \times 1$ column matrices $C_{j}^{\prime} \equiv U_{L}^{p \dagger} C_{j}$. Let

$$
U_{R}^{p}=\left(\begin{array}{c}
R_{1}  \tag{3.21}\\
R_{2} \\
R_{3}
\end{array}\right)
$$

where the $R_{1,2,3}$ are $1 \times 3$ row matrices. We know that

$$
\begin{equation*}
U_{L}^{p \dagger} M_{p}=\frac{1}{\sqrt{2}}\left(C_{1}^{\prime} v_{1}^{*}, C_{2}^{\prime} v_{2}^{*}, C_{3}^{\prime} v_{3}^{*}\right)=M_{u} U_{R}^{p \dagger}=\left(M_{u} R_{1}^{\dagger}, M_{u} R_{2}^{\dagger}, M_{u} R_{3}^{\dagger}\right) \tag{3.22}
\end{equation*}
$$

Therefore, $C_{j}^{\prime}=\left(\sqrt{2} / v_{j}^{*}\right) M_{u} R_{j}^{\dagger}$.
The couplings of the neutral scalars to the up-type quarks are given by

$$
\begin{align*}
- & \frac{1}{\sqrt{2}} \sum_{b=1}^{6} S_{b}^{0} \sum_{j=1}^{3} \mathcal{V}_{j b}^{*} \overline{p_{L}} \Delta_{j} p_{R}+\text { H.c. } \\
& =-\frac{1}{\sqrt{2}} \sum_{b=1}^{6} S_{b}^{0} \sum_{j=1}^{3} \mathcal{V}_{j b}^{*} \overline{p_{L}} C_{j} p_{R j}+\text { H.c. } \\
& =-\frac{1}{\sqrt{2}} \sum_{b=1}^{6} S_{b}^{0} \sum_{j=1}^{3} \mathcal{V}_{j b}^{*}\left(\overline{u_{L}}, \overline{c_{L}}, \overline{t_{L}}\right) U_{L}^{p \dagger} C_{j} R_{j}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right)+\text { H.c. } \\
& =-\frac{1}{\sqrt{2}} \sum_{b=1}^{6} S_{b}^{0} \sum_{j=1}^{3} \mathcal{V}_{j b}^{*}\left(\overline{u_{L}}, \overline{c_{L}}, \overline{t_{L}}\right) C_{j}^{\prime} R_{j}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right)+\text { H.c. } \\
& =-\sum_{b=1}^{6} S_{b}^{0} \sum_{j=1}^{3} \frac{\mathcal{V}_{j b}^{*}}{v_{j}^{*}}\left(\overline{u_{L}}, \overline{c_{L}}, \overline{t_{L}}\right) M_{u} R_{j}^{\dagger} R_{j}\left(\begin{array}{c}
u_{R} \\
c_{R} \\
t_{R}
\end{array}\right)+\text { H.c. } \tag{3.23}
\end{align*}
$$

Let us define $H_{j}=R_{j}^{\dagger} R_{j}$. The $H_{j}$ are three Hermitian matrices; since $U_{R}^{p}$ is unitary, $H_{1}+H_{2}+H_{3}=\mathbb{1}_{3 \times 3}$. The couplings of the neutral scalars to the top quark are given by

$$
\begin{equation*}
-m_{t} \sum_{b=1}^{6} S_{b}^{0} \sum_{j=1}^{3}\left(H_{j}\right)_{33} \bar{t}\left(\frac{\mathcal{V}_{j b}^{*}}{v_{j}^{*}} \gamma_{R}+\frac{\mathcal{V}_{j b}}{v_{j}} \gamma_{L}\right) t \tag{3.24}
\end{equation*}
$$

Thus, one given scalar $S_{b}^{0}$ couples to the top quark in the same way as the Higgs boson if

$$
\begin{equation*}
\sum_{j=1}^{3} \frac{v \mathcal{\nu}_{j b}\left(H_{j}\right)_{33}}{v_{j}}=1 \tag{3.25}
\end{equation*}
$$

For the action of $C P$ in the quark sector we choose

$$
C P:\left\{\begin{array}{l}
Q_{L} \rightarrow i \gamma_{0} C{\overline{Q_{L}}}^{T},  \tag{3.26}\\
n_{R} \rightarrow i \gamma_{0} C{\overline{n_{R}}}^{T}, \\
p_{R} \rightarrow i \gamma_{0} C S \bar{p}_{R}^{T} .
\end{array}\right.
$$

In this way,

$$
\begin{equation*}
C_{1}=C_{1}^{*}, \quad C_{2}=C_{3}^{*} . \tag{3.27}
\end{equation*}
$$

## 4 The scalar potential

The scalar potential may be separated into three pieces:

$$
\begin{equation*}
V=V_{\eta}+V_{\text {symmetric }}+V_{\text {soft }} . \tag{4.1}
\end{equation*}
$$

By definition, all the terms containing $\eta$ belong to $V_{\eta}$, whereas $V_{\text {symmetric }}+V_{\text {soft }}$ contains exclusively the $\phi_{j}$ with $j=1,2,3$. By definition, $V_{\eta}+V_{\text {symmetric }}$ is invariant under all the symmetries of the Lagrangian, i.e. under both $C P$ and the $\mathbb{Z}_{2}^{(j)}$ for $j=1,2,3$, whereas $V_{\text {soft }}$ breaks the $\mathbb{Z}_{2}^{(j)}$ softly but preserves $C P$. Obviously, in any term in $V_{\eta}+V_{\text {symmetric }}$ only even numbers of $\eta$ 's and of each of the $\phi_{j}$ 's can occur. We have

$$
\begin{align*}
V_{\eta}= & \mu_{\eta} \eta^{\dagger} \eta+\tilde{\lambda}_{1}\left(\eta^{\dagger} \eta\right)^{2} \\
& +\eta^{\dagger} \eta\left[\tilde{\lambda}_{2} \phi_{1}^{\dagger} \phi_{1}+\tilde{\lambda}_{3}\left(\phi_{2}^{\dagger} \phi_{2}+\phi_{3}^{\dagger} \phi_{3}\right)\right]+\tilde{\lambda}_{4} \eta^{\dagger} \phi_{1} \phi_{1}^{\dagger} \eta+\tilde{\lambda}_{5}\left(\eta^{\dagger} \phi_{2} \phi_{2}^{\dagger} \eta+\eta^{\dagger} \phi_{3} \phi_{3}^{\dagger} \eta\right) \\
& +\xi_{1}\left[\left(\phi_{1}^{\dagger} \eta\right)^{2}+\left(\eta^{\dagger} \phi_{1}\right)^{2}\right]+\left\{\xi_{2}\left[\left(\phi_{2}^{\dagger} \eta\right)^{2}+\left(\eta^{\dagger} \phi_{3}\right)^{2}\right]+\text { H.c. }\right\}, \tag{4.2}
\end{align*}
$$

where $\tilde{\lambda}_{1-5}$ and $\xi_{1}$ are real while $\xi_{2}$ is in general complex. We assume that the real coefficient $\mu_{\eta}$ is positive, so that the VEV $\left\langle\eta^{0}\right\rangle_{0}$ vanishes. Moreover, $\mu_{\eta}$ must be sufficiently larger than the Fermi scale squared, so that the terms with coefficients $\tilde{\lambda}_{2-5}$ cannot make $\mu_{\eta} \rightarrow \mu_{\eta}+\left(\tilde{\lambda}_{2}+\tilde{\lambda}_{4}\right)\left|v_{1}\right|^{2}+\left(\tilde{\lambda}_{3}+\tilde{\lambda}_{5}\right)\left(\left|v_{2}\right|^{2}+\left|v_{3}\right|^{2}\right)$ become negative.

Due to the symmetries, $V_{\text {symmetric }}$ is given by

$$
\begin{align*}
V_{\text {symmetric }}= & \mu_{1} \phi_{1}^{\dagger} \phi_{1}+\mu_{2}\left(\phi_{2}^{\dagger} \phi_{2}+\phi_{3}^{\dagger} \phi_{3}\right)+\lambda_{1}\left(\phi_{1}^{\dagger} \phi_{1}\right)^{2}+\lambda_{2}\left[\left(\phi_{2}^{\dagger} \phi_{2}\right)^{2}+\left(\phi_{3}^{\dagger} \phi_{3}\right)^{2}\right] \\
& +\lambda_{3} \phi_{1}^{\dagger} \phi_{1}\left(\phi_{2}^{\dagger} \phi_{2}+\phi_{3}^{\dagger} \phi_{3}\right)+\lambda_{4}\left(\phi_{1}^{\dagger} \phi_{2} \phi_{2}^{\dagger} \phi_{1}+\phi_{1}^{\dagger} \phi_{3} \phi_{3}^{\dagger} \phi_{1}\right) \\
& +\lambda_{5} \phi_{2}^{\dagger} \phi_{2} \phi_{3}^{\dagger} \phi_{3}+\lambda_{6} \phi_{2}^{\dagger} \phi_{3} \phi_{3}^{\dagger} \phi_{2}+\lambda_{7}\left(\phi_{2}^{\dagger} \phi_{3}\right)^{2}+\lambda_{7}^{*}\left(\phi_{3}^{\dagger} \phi_{2}\right)^{2} \\
& +\lambda_{8}\left[\left(\phi_{1}^{\dagger} \phi_{2}\right)^{2}+\left(\phi_{3}^{\dagger} \phi_{1}\right)^{2}\right]+\lambda_{8}^{*}\left[\left(\phi_{2}^{\dagger} \phi_{1}\right)^{2}+\left(\phi_{1}^{\dagger} \phi_{3}\right)^{2}\right] \tag{4.3}
\end{align*}
$$

with real $\lambda_{l}$ for $l=1, \ldots, 6$ and complex $\lambda_{7}$ and $\lambda_{8}$. The soft-breaking potential, which consists of terms of dimension two and abides by the $C P$ symmetry, is

$$
\begin{equation*}
V_{\text {soft }}=\mu_{3} \phi_{2}^{\dagger} \phi_{3}+\mu_{3}^{*} \phi_{3}^{\dagger} \phi_{2}+\mu_{4}\left(\phi_{1}^{\dagger} \phi_{2}+\phi_{3}^{\dagger} \phi_{1}\right)+\mu_{4}^{*}\left(\phi_{2}^{\dagger} \phi_{1}+\phi_{1}^{\dagger} \phi_{3}\right) . \tag{4.4}
\end{equation*}
$$

We firstly follow ref. [16] to investigate the minimum of $V_{\text {symmetric }}$. We write $\lambda_{l}=\left|\lambda_{l}\right| e^{i \alpha_{l}}$ for $l=7,8$ and

$$
\begin{equation*}
v_{1}=w_{1} e^{i \beta_{1}}, \quad v_{2}=w \sin \sigma e^{i \beta_{2}}, \quad v_{3}=w \cos \sigma . \tag{4.5}
\end{equation*}
$$

We require without loss of generality that $w_{1} \geq 0, w \geq 0$, and $\sigma$ is in the first quadrant. Let $F$ denote the sum of the terms of $\langle 0| V_{\text {symmetric }}|0\rangle$ that have a non-trival $\sigma$-dependence. One has

$$
\begin{align*}
4 F= & \lambda_{2} w^{4}\left(\sin ^{4} \sigma+\cos ^{4} \sigma\right)+\left(\lambda_{5}+\lambda_{6}\right) w^{4} \sin ^{2} \sigma \cos ^{2} \sigma  \tag{4.6a}\\
& +2\left|\lambda_{7}\right| w^{4} \sin ^{2} \sigma \cos ^{2} \sigma \cos \left(\alpha_{7}-2 \beta_{2}\right)  \tag{4.6b}\\
& +2\left|\lambda_{8}\right| w_{1}^{2} w^{2}\left[\sin ^{2} \sigma \cos \left(\alpha_{8}+2 \beta_{2}-2 \beta_{1}\right)+\cos ^{2} \sigma \cos \left(\alpha_{8}+2 \beta_{1}\right)\right] . \tag{4.6c}
\end{align*}
$$

In line (4.6a) we use $\sin ^{4} \sigma+\cos ^{4} \sigma=1-2 \sin ^{2} \sigma \cos ^{2} \sigma$. We require

$$
\begin{equation*}
-2 \lambda_{2}+\lambda_{5}+\lambda_{6}-2\left|\lambda_{7}\right|>0 . \tag{4.7}
\end{equation*}
$$

Then, the minimum of lines (4.6a) and (4.6b) is achieved when $\sin \sigma \cos \sigma=0$. This may also be the minimum of line (4.6c) because

$$
\begin{equation*}
\sin ^{2} \sigma \cos \left(\alpha_{8}+2 \beta_{2}-2 \beta_{1}\right)+\cos ^{2} \sigma \cos \left(\alpha_{8}+2 \beta_{1}\right) \geq-1 \tag{4.8}
\end{equation*}
$$

and the value -1 can always be obtained, irrespective of the value of $\sigma$, through suitable choices of $\beta_{1}$ and $\beta_{2}$. Thus, assuming $\sin \sigma=0$ instead of $\cos \sigma=0$, the minimum of $F$ is at $\sigma=0$, i.e. $v_{2}=0$, and $\alpha_{8}+2 \beta_{1}=\pi$. The latter relation, however, is irrelevant if $v_{1}=0$, because then $\beta_{1}$ is meaningless. We assume, indeed, that the coefficient $\mu_{1}$ is positive and so large that $v_{1}=0$. The minimum of $V_{\text {symmetric }}$ then has $v_{1}=v_{2}=0$.

In the limit $v_{1}=v_{2}=0$ it is easy to compute the scalar mass spectrum. Writing

$$
\begin{equation*}
\phi_{1}^{0}=e^{-i \alpha_{8} / 2} \frac{\rho_{1}+i \sigma_{1}}{\sqrt{2}}, \quad \phi_{2}^{0}=e^{i \alpha_{7} / 2} \frac{\rho_{2}+i \sigma_{2}}{\sqrt{2}}, \quad \phi_{3}^{0}=\frac{w+\rho_{3}+i \sigma_{3}}{\sqrt{2}} \tag{4.9}
\end{equation*}
$$

we find

$$
\begin{align*}
& m_{\phi_{1}^{+}}^{2}=\mu_{1}+\frac{\lambda_{3}}{2} w^{2},  \tag{4.10a}\\
& m_{\rho_{1}}^{2}=\mu_{1}+\left(\frac{\lambda_{3}+\lambda_{4}}{2}+\left|\lambda_{8}\right|\right) w^{2},  \tag{4.10b}\\
& m_{\sigma_{1}}^{2}=\mu_{1}+\left(\frac{\lambda_{3}+\lambda_{4}}{2}-\left|\lambda_{8}\right|\right) w^{2},  \tag{4.10c}\\
& m_{\phi_{2}^{+}}^{2}=\left(-\lambda_{2}+\frac{\lambda_{5}}{2}\right) w^{2},  \tag{4.10d}\\
& m_{\rho_{2}}^{2}=\left(-\lambda_{2}+\frac{\lambda_{5}+\lambda_{6}}{2}+\left|\lambda_{7}\right|\right) w^{2},  \tag{4.10e}\\
& m_{\sigma_{2}}^{2}=\left(-\lambda_{2}+\frac{\lambda_{5}+\lambda_{6}}{2}-\left|\lambda_{7}\right|\right) w^{2},  \tag{4.10f}\\
& m_{\rho_{3}}^{2}=2 \lambda_{2} w^{2} . \tag{4.10~g}
\end{align*}
$$

Moreover, $\sigma_{3}=G^{0}$ is the neutral Goldstone boson and $\phi_{3}^{+}$is the charged Goldstone boson, which are absorbed by the $Z^{0}$ and the $W^{+}$gauge bosons, respectively. The fields $\phi_{1}^{+}, \rho_{1}$, and $\sigma_{1}$ are heavy because of the large $\mu_{1}$. The scalar $\rho_{3}$ is to be identified with the Higgs boson. From $m_{\rho_{3}} \simeq 125 \mathrm{GeV}$ and $w \simeq 246 \mathrm{GeV}$ one obtains $\lambda_{2} \simeq 0.13$. The masses of $\phi_{2}^{+}, \rho_{2}$, and $\sigma_{2}$ cannot be very large if one wants to stay in the perturbative regime with respect to $\lambda_{5,6,7}$.

Now we take into account $V_{\text {soft }}$. This generates $v_{1,2} \neq 0$ due to the presence of terms linear in $\phi_{1}$ and $\phi_{2}$. For sufficiently small $v_{1}$ and $v_{2}$,

$$
\begin{align*}
& v_{1} \simeq-\frac{\mu_{4}^{*} v_{3}}{\mu_{1}+\frac{1}{2}\left(\lambda_{3}+\lambda_{4}\right)\left|v_{3}\right|^{2}+\lambda_{8}^{*} e^{-2 i \beta_{1}} v_{3}^{2}},  \tag{4.11a}\\
& v_{2} \simeq-\frac{\mu_{3} v_{3}}{\mu_{2}+\frac{1}{2}\left(\lambda_{5}+\lambda_{6}\right)\left|v_{3}\right|^{2}+\lambda_{7} e^{-2 i \beta_{1}} v_{3}^{2}} . \tag{4.11b}
\end{align*}
$$

In order to verify whether the scalar potential of equations (4.1)-(4.4) can produce a vacuum with the desired hierarchy of VEVs and yield acceptable scalar masses and couplings, we have performed a numerical scan of the parameter space of the potential. We have taken all the parameters of the scalar potential to be real, and we have also assumed real VEVs, parameterized as

$$
\begin{equation*}
v_{1}=v \cos \beta, \quad v_{2}= \pm \frac{v m_{\mu} \sin \beta}{\sqrt{m_{\mu}^{2}+m_{\tau}^{2}}}, \quad v_{3}=\frac{v m_{\tau} \sin \beta}{\sqrt{m_{\mu}^{2}+m_{\tau}^{2}}} \tag{4.12}
\end{equation*}
$$

where the angle $\beta$ may be either in the first or second quadrant. In this way we satisfy equation (2.7). It is desirable to have $\beta$ close to $\pi / 2$ so that $\left|v_{1}\right|$ is much smaller than $\left|v_{2,3}\right|$, because $m_{e} \propto v_{1}$, cf. equation (2.4); in our scan we have restricted $5 \leq|\tan \beta| \leq 400$.

Notice that equation (4.12) is used just as an Ansatz for our numerical study: nothing guarantees that the global minimum of the potential has real VEVs or, indeed, that it conserves the $\mathrm{U}(1)$ of electromagnetism. We also remind that, since in our model the $C P$ transformation effects $\phi_{2} \leftrightarrow \phi_{3}^{*}$, a vacuum with $v_{2} \neq v_{3}$ will in general lead to $C P$ violation even when the VEVs are real - indeed, we shall use equation (4.12) to fit for the observed $C P$ violation, cf. equations (4.17) below.

We have made the quartic couplings of the potential comply with certain basic restrictions for the model to make sense:

- The scalar potential has to be bounded from below (BFB), i.e. there should be no directions in field space along which the potential can tend to minus infinity. To find the BFB conditions one must study the behaviour of the scalar potential for specific directions along which the fields may tend to infinity and verify which combinations of parameters ensure that the potential is BFB. The set of necessary conditions ${ }^{8}$ that

[^5]we have enforced is (see refs. [21, 22])
\[

$$
\begin{array}{r}
\lambda_{1}>0 \\
\lambda_{2}>0 \\
L_{1} \equiv 2 \sqrt{\lambda_{1} \lambda_{2}}+\lambda_{3}+\left(\lambda_{4}-2\left|\lambda_{8}\right|\right) \Theta\left(2\left|\lambda_{8}\right|-\lambda_{4}\right)>0 \\
L_{2} \equiv 2 \lambda_{2}+\lambda_{5}+\left(\lambda_{6}-2\left|\lambda_{7}\right|\right) \Theta\left(2\left|\lambda_{7}\right|-\lambda_{6}\right)>0 \\
\sqrt{\lambda_{1}} L_{2}+2 \sqrt{\lambda_{2}} L_{1}-4 \lambda_{2} \sqrt{\lambda_{1}}+L_{1} \sqrt{L_{2}}>0 \tag{4.13e}
\end{array}
$$
\]

where $\Theta$ denotes the step function of Heaviside.

- The model must respect unitarity and perturbativity. Therefore, the quartic couplings of the potential cannot be arbitrarily large. We have imposed $\left|\lambda_{l}\right| \leq 20 \forall l=$ $1, \ldots, 8$; this should guarantee appropriate behaviour.
- The model has to obey the phenomenological constraint on the oblique parameter $T$, viz. $T=0.01 \pm 0.12$ [23]. The value of $T$ was computed through the formulae of ref. [19]. ${ }^{9}$

In the further discussion of this section, we use the following notation for the exact scalar mass eigenstates: charged scalars $H_{i}^{+}(i=1,2), C P$-even neutral scalars $h_{i}$ $(i=1,2,3)$, and pseudoscalars $A_{i}(i=1,2) .{ }^{10}$ There is the correspondence $H_{i}^{+} \leftrightarrow \phi_{i}^{+}$, $h_{i} \leftrightarrow \rho_{i}$, and $A_{i} \leftrightarrow \sigma_{i}$ between the exact and the approximate mass eigenstates, with the approximate masses given by equations (4.10). By definition, the mass of $A_{1}$ is larger than the mass of $A_{2}$ and the mass of $H_{1}^{+}$is larger than the mass of $H_{2}^{+}$; similarly, $m_{h_{1}}>m_{h_{2}}>m_{h_{3}}$.

With the above restrictions in place, we still have to implement in our numerical scan a scalar state $h_{3}$, corresponding to the Higgs boson, with mass $125 \pm 1 \mathrm{GeV}$ and almost "aligned", according to the discussion held in the previous sections, with the $\rho_{3}$ direction. Also, since the model does have FCNYI, it is very convenient that all the neutral scalars other than the lightest one be as heavy as possible. In our scan we have imposed a lower bound of 600 GeV on the masses of the charged scalars $H_{1,2}^{+}$, of the pseudoscalars $A_{1,2}$, and of the heavier $C P$-even scalars $h_{1,2} \cdot{ }^{11}$ Moreover, since the experimental constraints on FCNYI in the down-type-quark sector are much stronger than those in the up-type-quark sector, we have chosen to scan exclusively the extension 2 of our model to the quark sector, since that extension has no FCNYI in the down sector. Finally, in order to comply with current LHC experimental results [24], the 125 GeV -mass scalar $h_{3}$ must have couplings to the gauge bosons and to the heavy fermions close to the SM values. Specifically, in our scan we have demanded that:

[^6]

Figure 1. Scatter plot of the mass of the second heaviest $C P$-even scalar versus the absolute value of $v_{1}$.

- The coupling of $h_{3}$ to the gauge bosons be within $10 \%$ of its expected SM value, i.e.

$$
\begin{equation*}
g_{S V V}=\frac{v_{1} \mathcal{V}_{13}+v_{2} \mathcal{V}_{23}+v_{3} \mathcal{V}_{33}}{v} \geq 0.9 \tag{4.14}
\end{equation*}
$$

with the matrix $\mathcal{V}$ defined in equation (A.2). By definition, the third column of that matrix corresponds to the 125 GeV -mass neutral scalar $h_{3}$.

- The coupling of $h_{3}$ to the bottom quarks be within $10 \%$ of its expected SM value, i.e.

$$
\begin{equation*}
g_{S b b} \equiv \frac{v \mathcal{V}_{13}}{v_{1}}=1 \pm 0.1 \tag{4.15}
\end{equation*}
$$

- The coupling of $h_{3}$ to the tau leptons be within $10 \%$ of its expected SM value, i.e.

$$
\begin{equation*}
g_{S \tau \tau} \equiv \frac{v \mathcal{V}_{33}}{v_{3}}=1 \pm 0.1 \tag{4.16}
\end{equation*}
$$

In figure 1 we have plotted the mass of the second heaviest $C P$-even scalar, $h_{2}$, against the value of the VEV $v_{1}$. There are two features worth mentioning. Firstly, the value of $\left|v_{1}\right|$ may be very small, i.e. the value of $|\tan \beta|$ may be very large. Secondly, the mass of $h_{2}$ is never higher than 1.35 TeV . The first feature implies $\sin \beta \simeq 1$; therefore, the values of $v_{2}$ and $v_{3}$ in equations (4.12) are essentially constant: $\left|v_{2}\right| \simeq 14.6 \mathrm{GeV}$ and $v_{3} \simeq 245.3 \mathrm{GeV}$. The second feature arises from the need to keep the magnitudes of the quartic couplings in the perturbative regime, viz. $\left|\lambda_{i}\right| \leq 20$ for $i=1,3,4,5,6$. The other quartic couplings retain smaller magnitudes; we obtained $-12<\lambda_{7}<13$ and $-14<\lambda_{8}<2$ in our scan, while the coupling $\lambda_{2} \in[0.12,0.14]$ as predicted above.

Since we can find regions in the parameter space for which $\left|v_{1}\right| \ll\left|v_{2}\right| \ll\left|v_{3}\right|$, the expressions (4.10) constitute good approximations to the exact scalar masses. To illustrate this, in figure 2 we have plotted the exact mass of $H_{2}^{+}$against the approximate expression for that mass in equation (4.10d). As we can appreciate from the plot, the approximate formula describes quite reasonably the true value, though deviations $\lesssim 30 \%$ occur in some


Figure 2. Scatter plot of the approximate expression (4.10d) for the mass of the lightest charged scalar versus the true mass of that particle.


Figure 3. Scatter plot of the mass of the lightest charged scalar versus the mass of the lightest pseudoscalar.
cases. Similar results have been obtained for the approximate formulae for the masses of $h_{2}, A_{2}, h_{1}, A_{1}$, and $H_{1}^{+}$.

In figure 3 one observes that the mass of $A_{2}$ may be smaller than, but may also be as much as twice, the one of $H_{2}^{+}$. This is in spite of our enforcement of the experimental bound on the oblique parameter $T$, which might suggest the masses of $A_{2}, h_{2}$, and $H_{2}^{+}$to be almost degenerate; they are not. In figure 4 one observes the same as in figure 3 , but now for the heaviest charged scalar and the heaviest pseudoscalar. One sees once again that the masses of the heaviest scalars can differ considerably.

Comparing figures 3 and 4 one sees that the masses of the heaviest and the lightest scalars are not necessarily much different. This can be confirmed through figure 5 , where the masses of the two heavy neutral scalars are plotted against each other. One sees the $m_{h_{1}}$ and $m_{h_{2}}$ may be quite close to each other, whatever their average value.


Figure 4. Scatter plot of the mass of the heaviest charged scalar versus the mass of the heaviest pseudoscalar.


At this stage, we have shown that our model can reproduce a boson with mass roughly 125 GeV and couplings to the gauge bosons, to the bottom quarks, and to the tau leptons close to the expected SM values. The Yukawa-coupling matrix $\Gamma_{1}$, as described in section 3.2 , reproduces the known down-type-quark masses. We now have to show that the model can also reproduce both the up-type-quark masses and the CKM matrix. We take the values of the up-type-quark running masses, at the scale $m_{Z}$, from ref. [25] and the values of the CKM-matrix parameters from ref. [23]:

$$
\begin{align*}
m_{u} & =1.38_{-0.41}^{+0.42} \mathrm{MeV}, & m_{c} & =638_{-84}^{+43} \mathrm{MeV}, \quad m_{t}=172.1 \pm 1.2 \mathrm{GeV},  \tag{4.17a}\\
\left|V_{u s}\right| & =0.22536 \pm 0.00061, & \left|V_{c b}\right| & =0.04114 \pm 0.0012,  \tag{4.17b}\\
\left|V_{u b}\right| & =(355 \pm 15) \times 10^{-5}, & J & =306_{-20}^{+21} \times 10^{-7} . \tag{4.17c}
\end{align*}
$$

We have used the values of the parameters of the model that had previously been shown to obey all the constraints hitherto mentioned and we have searched for phenomenologically
acceptable values for the Yukawa couplings in equations (3.27):

$$
C_{1}=\left(\begin{array}{l}
f_{1}  \tag{4.18}\\
f_{2} \\
f_{3}
\end{array}\right), \quad C_{2}=\left(\begin{array}{c}
f_{4}+i f_{5} \\
f_{6}+i f_{7} \\
f_{8}+i f_{9}
\end{array}\right), \quad C_{3}=\left(\begin{array}{c}
f_{4}-i f_{5} \\
f_{6}-i f_{7} \\
f_{8}-i f_{9}
\end{array}\right),
$$

with real $f_{1}, \ldots, f_{9}$. Specifically, we have demanded in our fitting procedure that the up-type-quark mass matrix in equation (3.20), after being diagonalized as $V M_{p} M_{p}^{\dagger} V^{\dagger}=$ $\operatorname{diag}\left(m_{u}^{2}, m_{c}^{2}, m_{t}^{2}\right)$, gives both the right values for the masses of the up-type quarks and for the moduli of the matrix elements of the CKM matrix $V$. We moreover require that the coupling of $h_{3}$ to top quarks be within $10 \%$ of its SM value:

$$
\begin{equation*}
g_{S t t} \equiv \sum_{j=1}^{3} \frac{v v_{j} \mathcal{V}_{j 3}\left(V C_{j}^{\dagger} C_{j} V^{\dagger}\right)_{33}}{2 m_{t}^{2}}=1 \pm 0.1 \tag{4.19}
\end{equation*}
$$

We have also verified what constraints might arise from the limits on $b \rightarrow s \gamma$ observations. These would arise solely from the charged-scalars interactions, and should be similar in form to those found in 2HDM type-II, due to bottom and top quarks getting their masses from two different doublets. Due to the mixing of the two charged scalars, we have verified that the couplings of the lightest charged state to the fermions are suppressed compared to those one would obtain in a type-II 2HDM. As a result, the constraints from $b \rightarrow s \gamma$ obtained for our model are much less stringent than those found for the 2HDM type-II. Since we further consider only high masses for the charged scalars (above 600 GeV ), the constraints will be even less relevant.

We have found that, for each and every set of parameters of the scalar potential that has been used to produce figures $1-5$, it is possible to find values for $f_{1}, \ldots, f_{9}$ which lead to observables satisfying equations (4.17) and (4.19) almost perfectly. Indeed, most of the observables can be fitted at the $1 \sigma$ level, but at least one of the observables $m_{c},\left|V_{u b}\right|$, and $J$ can only be fitted at the $2 \sigma$ level. Thus, if all the observables except $m_{c}$ are within their $1 \sigma$ allowed domains, then the minimum pull ${ }^{12}$ of $m_{c}$ is 1.9 ; if all the observables except $\left|V_{u b}\right|$ have pull smaller than one in modulus, then $\left|V_{u b}\right|$ has a pull of at least 2.4; if all the observables but $J$ are within their $1 \sigma$ boundaries, then $J$ has a pull smaller than -1.8 . Altogether, the best fits that we were able to achieve have a value of $\chi^{2}$ - for the three quark masses and the four CKM-matrix observables in equations (4.17) - of 5.1.

An example of one of our best fits is provided in table 1.

## 5 Conclusions

In this paper we have shown that it is possible to unify the idea of a scotogenic neutrino mass model [13] with the enforcement of co-bimaximal lepton mixing. The latter is obtained via softly broken lepton numbers [1] and a non-standard $C P$ transformation which interchanges the $\mu$ and $\tau$ flavours. Such a $C P$ transformation procures $\theta_{23}=45^{\circ}$ and $\delta= \pm \pi / 2$ in the

[^7]| parameter | value | observable | value |
| :---: | :---: | :---: | :---: |
| $\beta(\mathrm{rad})$ | 1.5732963215865827 | $m_{h_{3}}(\mathrm{GeV})$ | 125.0 |
| $\lambda_{1}$ | 17.135112092706517 | $m_{h_{2}}(\mathrm{GeV})$ | 739.2 |
| $\lambda_{2}$ | 0.13092447205288404 | $m_{h_{1}}(\mathrm{GeV})$ | 951.7 |
| $\lambda_{3}$ | 15.624853371379327 | $m_{A_{2}}(\mathrm{GeV})$ | 1106 |
| $\lambda_{4}$ | -11.846787927249578 | $m_{A_{1}}(\mathrm{GeV})$ | 1281 |
| $\lambda_{5}$ | 19.99999999813406 | $m_{H_{2}^{+}}(\mathrm{GeV})$ | 773.0 |
| $\lambda_{6}$ | 16.363914697200098 | $m_{H_{1}^{+}}(\mathrm{GeV})$ | 1193 |
| $\lambda_{7}$ | -9.030984509839026 | $g_{S V V}$ | 0.9925 |
| $\lambda_{8}$ | -2.6314236783145977 | $g_{S b b}$ | 1.000 |
| $\mu_{4}\left(\mathrm{GeV}{ }^{2}\right)$ | 2140.7424941612453 | $g_{S \tau \tau}$ | 1.000 |
| $f_{1}$ | 0.0024483113150037543 | $T$ | 0.01000 |
| $f_{2}$ | -0.03085374618190331 | $m_{u}(\mathrm{MeV})$ | 1.665 |
| $f_{3}$ | -0.20462462612388946 | $m_{c}(\mathrm{MeV})$ | 679.1 |
| $f_{4}$ | 0.0032554425959401188 | $m_{t}(\mathrm{GeV})$ | 170.9 |
| $f_{5}$ | 0.00697096592835829 | $\left\|V_{u s}\right\|$ | 0.2259 |
| $f_{6}$ | -0.001672825126610988 | $\left\|V_{c b}\right\|$ | 0.04144 |
| $f_{7}$ | -0.040296075343726166 | $\left\|V_{u b}\right\|$ | 0.003694 |
| $f_{8}$ | 0.7561059024727611 | $J$ | 0.00002706 |
| $f_{9}$ | -0.6259667363570083 | $g_{S t t}$ | 0.9926 |
|  |  |  |  |

Table 1. The values of the parameters and of the observables for one of our fits. The sign of $v_{2}$ - see equation (4.12) - is positive for this fit. The values of $\mu_{1,2,3}$ were computed by using the stationarity equations for the vacuum.
lepton mixing matrix, while $\theta_{13}$ remains undetermined; this is in good agreement with the data. In a scotogenic model, the neutrino masses are generated through a one-loop diagram involving the dark sector, which consists of right-handed neutrinos and a scalar gauge doublet $\eta$ which has zero VEV. Thus, a scotogenic model combines neutrino-mass suppression through the seesaw mechanism and through radiative mass generation.

Our model contains three scalar doublets with nonzero VEVs. Therefore, we wanted to demonstrate that a scalar $h_{3}$ with mass 125 GeV can be accommodated in our model. We have shown that this scalar can be made to have couplings to the gauge bosons and to the heavy fermions very close to those of the Higgs particle. Since the non-standard $C P$ transformation interchanges, besides the $\mu$ and $\tau$ flavours, also two of the scalar doublets, it is non-trivial to make all the scalars other than $h_{3}$ heavy. Still, we have found that all of them can be made to have masses above 600 GeV .

We have also demonstrated that the symmetries of our model may consistently be extended to the quark sector, correctly reproducing all the quark masses and the CKM matrix. There are neutral scalar-mediated flavour changing currents; however, it is possible to choose the model's symmetries so that they occur only in the up-type-quark sector, for which the experimental constraints on such currents are much looser. The fit to the quark sector is at the $2 \sigma$ level, but with many observables falling within their $1 \sigma$ uncertainty intervals.

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## A One-loop neutrino mass corrections

In this appendix we collect some formulae from ref. [12], adapting them to the model in this paper. In particular, we set $n_{L}=n_{R}=3$ and $n_{H}=4$, where $n_{L}, n_{R}$, and $n_{H}$ are, in the notation of ref. [12], the numbers of fermion families, of right-handed neutrino singlets, and of scalar doublets, respectively. The Yukawa Lagrangian of the right-handed neutrinos in equation (1) of ref. [12] is given by

$$
\begin{equation*}
\mathcal{L}_{\nu_{R} \text { Yukawa }}=-\overline{\nu_{R}}\left(\sum_{k=1}^{4} \tilde{\phi}_{k}^{\dagger} \Delta_{k}\right) D_{L}+\text { H.c. } \tag{A.1}
\end{equation*}
$$

The notation for the physical neutral scalars is best explained in ref. [19]. The neutral component of the scalar doublet $\phi_{k}(k=1,2,3,4)$ has VEV $\langle 0| \phi_{k}^{0}|0\rangle=v_{k} / \sqrt{2}$ and is written as

$$
\begin{equation*}
\phi_{k}^{0}=\frac{1}{\sqrt{2}}\left(v_{k}+\sum_{b=1}^{8} \mathcal{V}_{k b} S_{b}^{0}\right), \tag{A.2}
\end{equation*}
$$

where the complex matrix $\mathcal{V}$ is $4 \times 8$. The neutral Goldstone boson is $S_{1}^{0}$ and the remaining seven $S_{b}^{0}$, for $b=2, \ldots, 8$, are physical neutral scalars with masses $m_{b}$. For each physical neutral scalar we define, following ref. [12], the matrix

$$
\begin{equation*}
\Delta_{b} \equiv \sum_{k=1}^{4} \mathcal{V}_{k b} \Delta_{k} \tag{A.3}
\end{equation*}
$$

Then, the final result in equation (53) of ref. [12] is

$$
\begin{align*}
\delta \mathcal{M}_{\nu}= & \sum_{b=2}^{8} \frac{m_{b}^{2}}{32 \pi^{2}} \Delta_{b}^{T} W^{*}\left(\frac{1}{\widetilde{m}} \ln \frac{\widetilde{m}^{2}}{m_{b}^{2}}\right) W^{\dagger} \Delta_{b}  \tag{A.4a}\\
& +\frac{3 g^{2}}{64 \pi^{2} c_{w}^{2}} M_{D}^{T} W^{*}\left(\frac{1}{\widetilde{m}} \ln \frac{\widetilde{m}^{2}}{m_{Z}^{2}}\right) W^{\dagger} M_{D} . \tag{A.4b}
\end{align*}
$$

The sum in line (A.4a) includes only the physical neutral scalars. Line (A.4b) includes the contributions from the loop with a $Z^{0}$ and from the loop with a neutral Goldstone boson. In that line, $M_{D}=\sum_{k=1}^{4}\left(v_{k} / \sqrt{2}\right) \Delta_{k}$ is the Dirac neutrino mass matrix; in the model in this paper that matrix vanishes, because both the matrices $\Delta_{1,2,3}$ and the VEV $v_{4}$ are null. Therefore, for this paper only line (A.4a) matters.

In equation (A.4), the $3 \times 3$ unitary matrix $W$ is the one that diagonalizes $M_{R}$ according to equation (51) of ref. [12]:

$$
\begin{equation*}
W^{\dagger} M_{R} W^{*}=\widetilde{m} \equiv \operatorname{diag}\left(m_{4}, m_{5}, m_{6}\right) \tag{A.5}
\end{equation*}
$$

where $m_{4,5,6}$ are the masses of the physical heavy neutrinos.

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[^0]:    ${ }^{1} \mathrm{~A}$ different way for obtaining co-bimaximal mixing, not involving the condition (1.1), has been recently proposed in refs. [7, 8].

[^1]:    ${ }^{2}$ Recently, another such model, but which employes a completely different mechanism, has been proposed in ref. [14]. The model of ref. [14] is more complicated than the one presented in this paper for several reasons: (1) it has two types of dark matter, one of them protected by a $U(1)$ symmetry and the other one by a $\mathbb{Z}_{2}$ symmetry. (2) It has several more fields in the dark sector. (3) The masses of the charged leptons are of radiative origin, just as those of the neutrinos. (4) The soft breaking of the symmetries occurs in two steps, with an $A_{4}$ symmetry in the dimension-four terms being softly broken to $\mathbb{Z}_{3}$ through dimension-three terms and that $\mathbb{Z}_{3}$ being softly broken through dimension-two terms.
    ${ }^{3}$ Such scalar doublets have been dubbed 'inert' in ref. [15].

[^2]:    ${ }^{4}$ We recall that in a renormalizable theory a symmetry is said to be broken softly when all the symmetrybreaking terms have dimension smaller than four. This leaves open two possibilities: either they have both dimension two and dimension three or they have only dimension two. Soft symmetry breaking is consistent in quantum field-theoretic terms because, when using it, the dimension-four symmetry-violating terms generated by loops are finite. The soft breaking of (super)symmetries is extensively used in model-building; in particular, all supersymmetric models contain soft supersymmetry-breaking terms.

[^3]:    ${ }^{5}$ We might accept the soft breaking of $C P$ by quadratic terms in the scalar potential; that soft breaking by terms of dimension two would not disturb the dimension-three terms in $\mathcal{L}_{\text {Majorana }}$. But, for the sake of simplicity, we shall refrain in this paper from such a soft breaking.
    ${ }^{6}$ Ours is a model of 'real $C P$ violation', i.e. $C P$ violation originates in the inequality of two VEVs, even if those VEVs are real $[17,18]$.

[^4]:    ${ }^{7}$ Notice that the quark mixing matrix is $V=\sum_{j=1}^{3} \hat{R}_{j}^{\dagger} \hat{R}_{j}$.

[^5]:    ${ }^{8}$ A set of necessary and sufficient BFB conditions was obtained for the two-Higgs-doublet model in ref. [20], but the procedure described therein cannot be generalized to models with a larger scalar content.

[^6]:    ${ }^{9}$ We have explicitly checked that the bounds on the oblique parameter $S$ never give additional restrictions to this model.
    ${ }^{10}$ Note that, since in our fit we have assumed both the parameters of the potential and the VEVs to be real, the scalar sector of the model conserves $C P$, hence there are well-defined neutral scalars and pseudoscalars.
    ${ }^{11}$ We have also imposed an upper bound of 1500 GeV on all the scalar masses.

[^7]:    ${ }^{12}$ As usual, we define the 'pull' of an observable as the difference between its fitted value and its mean value divided by the standard deviation.

